The GLAST Silicon-Strip Tracker Detectors and Electronics

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- Introduction to GLAST.
- GLAST Tracker Electronics Requirements.
- Detectors.
- Amplifiers and Discriminators.
- 1997 Beam Test.
- Readout Architecture and ASIC development.
- Mechanical Issues.
- Conclusions.



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- We detect gamma-rays in the energy range 10 MeV to tens of GeV by tracking and measuring the pair-conversion electrons.
- This energy range is inaccessible • from the ground.

kull sky view of the universe, from

The present EGRET experiment on the CGRO satellite has revolutionized this field of astrophysics but has nearly run out of *spark-chamber* gas.

Compton Gamma Ray Observatory

December, 1997 International Symposium on Semiconductor Tracking Detectors



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The Next Generation

- The use of 1990's tracking technology can produce factors of 10 to 100 improvement over EGRET's sensitivity while still keeping within a moderate budget.
- Self-triggering tracking allows one to dispense with the time-of-flight system—the resulting flat instrument provides a nearly fullsky field of view.

GLAST will provide nearly continuous monitoring of the characteristically transient sources of the gamma-ray sky.





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Why Silicon-Strip Detectors?

- Mature, proven, reliable, and robust technology.
- The ultimate vertex-detector resolution is not needed, but the relatively fine pitch provides excellent tracking for optimal
 - angular resolution,
 - 2-track separation,
 - and background rejection.
- High efficiency, with very low noise occupancy.
- No consumables, such as gas.
- Relatively low voltage.
- Fast—negligible dead time, self triggering.
- Cost—small part of the overall mission price, even for 100 sq m.

Alternative tracking technologies proposed by other groups:

- Gas Microstrips:
 - Tall, narrow cells track efficiently only at 90° incidence, at best.
 - Not proven to be a reliable technology (crucial in space!).
- Scintillating fibers:
 - Photoelectron yield is marginal, even with 1 mm fibers.
 - Only viable readout system (image intensifier + CCD) is unwieldy, slow, and needs very high voltage.





GLAST is an assembly of identical modules, each with a plastic scintillator veto shield, a silicon-strip tracker, and a cesium iodide calorimeter.

Each tracker module has 16 tracking layers, each with an x, y pair of singlesided silicon-strip detector planes. Each of the top 13 layers has a ~3.5% lead converter foil.

The current tracker baseline design calls for 32 cm square detector planes, each with a 5×5 array of 6.4cm detectors.



GLAST Baseline Conceptual Design



Electronics Requirements & Goals

Challenge: 1.3 million readout channels operating with high reliability in a space environment.

- Power less than $\approx 250 \,\mu$ W/channel, including amplifiers and digital readout. •
- Low noise occupancy (<0.05%) and good threshold uniformity. ۲
 - Expected detector loading is about 1.2 pF/cm \times 32 cm = 38 pF.
 - Detector thickness is 400 μ m, for about 5.3 fC/MIP.
- Microsecond peaking time for the amplifiers. ۲
- Sufficient redundancy to be immune to single-point failures. ۲
- Self triggering.
- Radiation hard to 10 kRad with latch-up immunity. •
- <1% dead-time at a 10 kHz trigger rate. •
- Sparse readout and data formatting close to the front end.

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GLAST Detectors

1997 Beam Test

- Hamamatsu Photonics
- AC coupled.
- Punch-through biasing.
- 236 µm strip pitch.
- 57 µm strip width.
- 500 µm thick.

Some measured properties:

Full depletion: ~140 V. Punch-through voltage: ~9 V. Strip Capacitance: ~1.2 pF/cm. Radiation damage: ~15 nA/strip after 10 kRad dose at 20°C.

Prototype Tower Module

(under construction)

- Hamamatsu Photonics
- AC coupled.
- Polysilicon biasing (30-80 M Ω).
- 194 µm strip pitch.
- 50 µm strip width.
- 400 µm thick.

Five detectors will be connected in series, to give 32-cm strips, each with ~38 pF capacitance.





The expected radiation dose for GLAST is relatively small (5 to 10 kRad), but measured effects on our prototype detectors indicate that some caution is needed.

- The new detector design employs polysilicon bias resistors, • rather than punch-through structures.
- The amplifier time constant will be kept below about $1.5 \,\mu s$.



International Symposium on Semiconductor Tracking Detectors December, 1997





Channel



The measured noise is about 25% greater than what is naïvely expected from shot noise due to the measured 15 nA leakage current.

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Noise measured from

channels.

threshold curves on 10

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Amplifiers and Discriminators

1500

1300

300

100

ENC=174 + 32 C

External Capacitors

16-Channel Prototype

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- Gain (shaper output): $\approx 150 \text{ mV/fC}$.
- Peaking time: $\approx 1.6 \,\mu s$. ٠
- Power consumption, including bias circuitry: 150 µW/channel
- Noise: ENC= $174 + 32 \times C$ electrons, with C in pF.
- RMS threshold variation: typically 6 to 10 mV across a single chip.



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Preamplifier Design

• Standard folded cascode amplifier with 2V bias for the front end, to save power.

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- ≈25 µA bias current set by an external resistor.
- Slow differential amplifier stabilizes the bias point and provides a continuous reset.
- Input impedance ≈5 kΩ gives
 ≈200 ns time constant with GLAST 38 pF detector load.
- Open loop gain: 64 dB at 0 Hz
- Power: ≈90 µW



Preamplifier schematic.



4uA



Shaping Amplifier Design

30nA (

• AC coupling from the preamplifier.

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- Conventional cascode amplifier with capacitive feedback.
- Slow differential amplifier in the feedback provides the "differentiation" function *and* stabilizes the output bias point. (Ref.: I. Kipnis, LBNL)
- Open loop gain: 62 dB at 0 Hz
- Voltage gain: ≈26
- Peaking time: $\approx 1.3 \,\mu s$
- Pulse shape: reset current source makes a tail that is more linear than exponential, except at the lowest pulse heights.
- Power: $\approx 35 \,\mu W$



Shaping Amplifier Schematic

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October 1997 Beam Test

- 12 planes of detectors (6*x* and 6*y*) read out by 72 32channel 2nd generation amplifier-discriminator chips. (Plus CsI calorimeter & anti-coinc. syst.)
- Interchangeable Pb converter foils of various thicknesses.
- One month of running in electron and tagged photon beams at SLAC.





Four double-sided modules and two Pb converter foils.









Several runs were taken at various detector bias voltages and various discriminator thresholds. With full depletion (V \geq 140 V), the efficiency is essentially 100% even well above the 1.5 fC threshold used during normal running.

- High energy electrons.
- CsI calorimeter was used to reject events with more than one electron.
- Normal incidence.
- Only detector planes with no dead or noisy channels were included.
- Tracks were required to have hits in planes both before and after the plane under test.



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Readout Architecture



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• 25 readout chips per layer.

- Two redundant digital readout-controller chips on each layer, one at each end.
- Data and Fast-OR (for trigger) move from chip to chip in either direction.
- Any readout chip can be reprogrammed and controlled by either controller chip, so if any single chip is bad, all others can still be read.
- All communication and control is via serial lines.
- Formatted hit lists are read from the controller chips in a token-controlled daisy chain.

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Front-End Readout Chip

- The 64-channel chip submitted this week to MOSIS includes the following features:
 - Calibration mask, to select any subset of channels to be pulsed.
 - Two 7 bit DACs, for setting calibration and threshold levels.
 - Separate masks for data and trigger output (Fast-OR).
 - 8-deep FIFO event buffer.
 - Dual redundant serial command decoders.
 - Dual redundant output shift registers and trigger outputs.
 - Bypasses of output registers, to avoid clocking out data from empty chips.
 - External communication via lowvoltage-swing differential signals.









- Control initialization, calibration, and readout of the front-end readout chips.
- Sparse readout—build list of hitstrip addresses.
- Calculate the time-over-threshold of the prompt trigger output.
- Build events and coordinate the readout with neighboring layers via a serial data line and a token.
- External communication via low-voltage-swing differential lines.
- Currently being designed for the HP 0.8 µm process using the CMOSX standard cells. The logic design is synthesized from Verilog code.



Simplified block diagram of the readout controller chip.

Mechanical-Thermal Issues

The GLAST Silicon-Strip Tracker

- Rigid, low-mass supports for the • detector planes \Rightarrow composite "trays."
- Minimum gap between trays, to keep • the *x*, *y* detector pairs as close as possible to the converter foils.
- Electronics on the sides of the trays, • to minimize dead area.
- Flex circuits to connect the detectors • to the electronics.
- Passive cooling via the tower walls.

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Two trays, oriented 90° with respect to each other. Detectors on top and bottom.

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- Glued to the surface of the tray.
- Contacts the back of each of 25 detectors at 3 points to supply the bias potential.
- Copper shield layer connected to ground to prevent noise pickup from the tray structure.
- 1610 fine gold traces on one edge, to carry signals, bias, and ground around the corner of the tray to the electronics.



inductance paths to the decoupling capacitors.



Hybrid Circuits

- The hybrid circuits include
 - bypass capacitors for each readout chip and the detector bias,
 - resistors to set the front-end bias current, and termination resistors for the digital signal lines.
 - Power supply filtering and fuses.

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- Temperature monitors.
- Eight layers are used in the layout, including entire planes for power, ground, analog 2V, and detector bias. The central type has already been fabricated.





Engineering Prototype Tower

- A single GLAST tower is now being built:
 - complete by early 1999,
 - verify engineering and develop assembly procedures,
 - subject to environmental testing,
 - beam test in electron, photon, and hadron beams,
 - and possibly fly in a balloon.
- Current NASA strategic plan:
 - New start for construction of GLAST in FY 2000. All R&D must be finished by then!
 - Launch in mid 2004.







During the next decade, siliconstrip technology will be greatly expanding our knowledge of the surprisingly many enigmatic gamma-ray sources in the h<u>eavens:</u>

- galactic core
- pulsars
- supernova remnants
- active galactic nuclei





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GLAST Collaboration



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Naval Research Laboratory, SACLAY-France, Sonoma State University, Stanford University: HEPL & SLAC, University of California Santa Cruz, University of Chicago, University of Rome- Italy, University of Tokyo-Japan, University of Washington

10 Astronomy/Astrophysics (black) + 10 HEP (red) Institutions (About 90 Collaborators)