

## Goal

What Can HEP and Astrophysics Practice Teach Each Other
Astrophysics:
aims at simple formulae (very fast)
calculates Sigmas directly
hope it's a good formula
HEP:
calculates probabilities by MC (general; slow) translates into Sigmas for communication loses track of analytic structure

Report at PHYSTAT2003 (Sept)

## Executive (Jordanian) Summary

- For high and moderate Non, Noff, Li Ma Eq 17 fine
- Anything works for Crab, but not for short GRB's
- Li Ma Eq 9 not too bad
- Bad formula typically overestimates significance
- Probably should use Binomial Test for small N
- Optimal Frequentist, and Plausible Bayesian, Technique
- want some MC confirmation
- numerically, more work than Li Ma Eq 9
- Interesting relations exist among methods
- Bayes with Gamma (not Gaussian) = Binomial
- And same as Alexandreas et. al. (possibly within a constant factor)...
- Li MaEq 9 = Binomial for large N
- Fraser-Reid Approximation Promising but not done


## Significance

- Z value: ~ $\operatorname{Normal}(0,1)$
(Milagro; Li Ma)
- The art is to pick a good variable for this


## More Generally:

- P(more extreme "signal" | background)
- Assume Null Hypothesis: background only
- Translate probability p into Z value by

$$
Z=\Phi^{-1}(p) ; \quad \Phi=\int_{-\infty}^{Z} e^{-t^{2} / 2} d t
$$

$$
Z \approx \sqrt{u-\operatorname{Ln} u} \approx \sqrt{-\operatorname{Ln} p}, \quad u=-2 \operatorname{Ln}(p \sqrt{2 \pi})
$$

## Prospective vs. Observed Significance

- This discussion: Observed Significance
- Post-hoc: (after data)
- Prospective Observability (before data) involves more :
- definition of $Z$, as for post-hoc; but also:
- Choice of Zmin $\quad=\max \mathrm{P}$ (observed|background)
- Very Similar to post-hoc: Zmin to make observation claim
- Consideration of probability of meeting criterion
- Simplest calculation:
- Non $=\mu_{\text {on }}+\alpha \mu_{\text {off }} ; \operatorname{Noff}=\mu_{\text {off }} \quad$ (ignores fluctuations)
- Significance for Expected Conditions
- Optimistic: crudely, $1 / 2$ time less signal; or $1 / 2$ time more background!
- Better: Source Strength for $50 \%$ probability of observation? $90 \%$ ?
- More related to Lazar's "upper bound" criterion
- Similar discussions in HEP literature


## Backgrounds in Astro and HEP

- Astrophysics: on-source vs. off-source
- side observation with $\alpha=$ Ton/Toff (sensitivity ratio)
$-\mathrm{b}=\alpha$ Noff; $\mathrm{db}=\alpha \sqrt{ }$ Noff
$-\alpha=(\mathrm{db})^{2} / \mathrm{b} \quad$ (deduced from above)
- HEP: estimate background in defined signal region
- Sometimes a sideband measurement, like Astrophysics
- Often a MC estimate; rescaled to signal sensitivity
- More often a sum of terms of both types
$-\mathrm{b} \pm \mathrm{db} \quad \mathrm{db}$ : uncertainties in quadrature
$-\boldsymbol{\alpha}=(\mathbf{d b})^{\mathbf{2}} / \mathbf{b} \quad$ I'll use as a definition of effective $\alpha$ Can apply astrophysics formulae


## Li and Ma Equations <br> $$
\mathrm{Z}=\mathrm{S} / \sigma(\mathrm{S})
$$ <br> $$
\mathrm{S}=\text { Non }-\mathrm{b} \quad \mathrm{~b}=\alpha \text { Noff }
$$

N is observation; $\quad \mathrm{b}$ is background estimate
Eq 5: $\operatorname{Var}(S)=\operatorname{Var}(\operatorname{Non})+\operatorname{Var}(\mathrm{b})=\operatorname{Non}+\alpha^{2}$ Noff
Ignores key null hyp constraint: $\mu_{\text {on }}=\alpha \mu_{\text {off }}$ (anti-signal bias!)
Eq 9: $\operatorname{Var}(S)=\alpha($ Non + Noff $)$
Obeys constraint; uses Non and Noff to estimate $\mu_{\text {off }}$
Eq 17: Log Likelihood Ratio (Wilks’ Theorem)

$$
\begin{gathered}
Z=\sqrt{2} \sqrt{x \bullet \operatorname{Ln}\left[\left(\frac{1+\alpha}{\alpha}\right)\left(\frac{x}{x+y}\right)\right]+y \bullet \operatorname{Ln}\left[(1+\alpha)\left(\frac{y}{x+y}\right)\right]} \\
\mathrm{x}=\text { Noff; } \mathrm{y}=\operatorname{Noff}
\end{gathered}
$$

## Li and Ma Variant

- Apply null by using only Noff to estimate $\operatorname{Var}(\mathrm{Non})$ and $\operatorname{Var}(\mathrm{b})$
- Obviously, bad if $\alpha>1$
- Eq 5c: $\operatorname{Var}(S)=\alpha(1+\alpha)$ Noff
- Get Eq 9 if use both (Max Likelihood)


## Other Frequentist Methods

## Ignoring uncertainty in b:

- $\mathrm{S} / \sqrt{ } \mathrm{b}$ Li Ma 10a
- Poisson( $\geq$ Non $\mid b)$
- Feldman \& Cousins?
(often much better) confidence limits!
- For significance, just Poisson $(\geq k \mid b)$, I believe


## Using Uncertainty in b :

- $b+d b$ instead of $b$ in above (I've seen it!)
- Near-Constant Variance (Zhang and Ramsden)

$$
Z_{V S}=\frac{2}{\sqrt{1+\alpha}}\left(\sqrt{x+\frac{3}{8}}-\sqrt{y+\frac{3}{8}}\right)
$$

- Fraser Reid
- Binomial Test


## Fraser and Reid

- Interesting approximate method (last 15 yr )
- Significance from likelihood curve
- Combine Z(Likelihood Ratio), Z(t/б)
- correct each other to $\mathrm{O}\left(\mathrm{n}^{-1.5}\right)$
- One version: improved $Z$ value
- Redo algebra for each new kind of problem
- I'm still working to apply it to Non, Noff fully
- Fast \& simple numerically to apply formula


## Binomial Test

## For Ratio of Poisson Means (Compare means for on, off measurements)

- UMPU (Uniformly Most Powerful Unbiased)
- If best test, probably it's using the best variable
- Holds $\mathrm{k}=$ Non+Noff fixed (nuisance parameter)
- Test is $\operatorname{PrBinomial}(\geq \mathrm{Non} \mid \mathrm{p}, \mathrm{k}), \mathrm{p}=\alpha /(1+\alpha)$
- Not in common use; probably should be Known in HEP and Astrophysics: not as optimal, nor standard procedure
- Zhang and Ramsden claim too conservative for $Z$ small? Even if true, we want $Z>4$

Experimental Astronomy 1 (1990) 145-163; I have pdf

- Closed form in term of special functions, or sums
- Applying for large N requires some delicacy; slower than Eq 17!
- Gaussian Limit of Binomial Test is Li Ma Eq 9!


## Bayesian Methods

- Allow for correlations among background contributions (MC integration)
- Extension to efficiency, upper limits natural
- In common use in HEP
- Cousins \& Highland "smeared likelihood" efficiency
- Predictive Posterior (after background measurement)
- Natural avenue for connection with p-values

But: typical Bayes analysis isn't significance, but odds ratio

- Truncated Gaussian often used to represent db
- A flat prior for background, gives gamma for db
- P value calc using gamma: (same(?) as Alexandreas)
- same as Frequentist Binomial Test


## Comparing the Methods

- Some test cases from literature
- Range of Non, Noff values
- Different $\alpha$ values
- Color Code Accuracy
- Assume Frequentist Binomial as Gold Standard
- May change after I've run Monte Carlo

|  | Top 11 | Top 2 | Top 3 | Crab $12>2.5$ | Crab 3 5 | Whipple | Hegria | Alexandr | Fake | Zhang 1 | Zhang2 |
| :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: | :---: |
| Non | 9 | 17 | 6 | 2,119,449 | 167,599 | 498,426 | 523 | 4 | 200 | 67 | 50 |
| Noff | 17.83 | 40.11 | 1878 | 23,671,193 | 1,604,910 | 493,434 | 2327 | 5 | 10 | 15 | 55 |
| alpha | 02132 | 0.0947 | 00692 | 0.0091 | 0.0991 | 1.000 | 0.167 | 0.2 | 10.0 | 2.0 | 0.5 |
| Nb | 38 | 3.6 | 1.3 | 2,109,732 | 166,213 | 493,434 | 308.6 | 1.0 | 100.0 | 30.0 | 27.5 |
| Ne - Non = Nb | 52 | 13.2 | 4.7 | 9717 | 1376 | 4992 | 134.4 | 3.0 | 100 | 37 | 22.5 |
| SigmatiNb) | 09 | 0.6 | 0.3 | 433.6 | 121.7 | 702.4 | 6.1 | 0.45 | 31.6 | 7.75 | 371 |
| Sigmant | 0237 | 0.158 | 0.231 | 0.000206 | 0.000732 | 0.00142 | 000207 | 0.447 | 0.316 | 0.258 | 0.136 |
| Reported a | 2700.02 | 2.00E-06 | 300E-03 |  |  |  |  |  |  |  |  |
| Reported S | 1.9 | 4.6 | 2.7 | 6.4 | 32 | 50 | 5.9 |  |  | 3 | 3 |
|  |  |  |  |  |  |  |  |  |  |  |  |
| Li Ma: |  |  |  |  |  |  |  |  |  |  |  |
| Eq5 | 1.68 | 3.17 | 1.90 | 6.397 | 3.22 | 5.01 | 5.54 | 1.46 | 289 | 328 | 282 |
| Eq9 | 2.17 | 567 | 3.59 | 6.409 | 3.23 | 5.01 | 616 | 224 | 218 | 289 | 311 |
| Eq 17 | 1.99 | 4.57 | 2.81 | 6.405 | 3.23 | 5.01 | 593 | 1.95 | 236 | 304 | 302 |
| Eq 50 | 2.42 | 6.47 | 399 | 6.410 | 3.23 | 5.03 | 631 | 274 | 302 | 390 | 350 |
| E4 10a - SGqu日 | 2.67 | 6.77 | 4.12 | 6.69 | 3.38 | 7.11 | 682 | 3.00 | 10.00 | 6.76 | 4.29 |
| $\mathrm{E}_{4} 10 \mathrm{Ve}+1 \mathrm{sgma}$ | 2.40 | 629 | 372 | 669 | 3.37 | 7.10 | 675 | 2.49 | 872 | 6.02 | 4.09 |
|  |  |  |  |  |  |  |  |  |  |  |  |
| Poisson | 2.14 | 487 | 2.34 | 6.69 | 3.37 | 7.09 | 6.44 | 206 | 7.72 | 5.76 | 3.80 |
| Puisent 4 Isigma | 1.64 | 4.47 | 2.46 | 6.336 | 3.07 | 6.09 | 601 | 1.56 | 551 | 4.24 | 304 |
| Binomial, Bayes prex | 1.82 | 4.46 | 2.63 | 6.4048 | 3.23 | 5.01 | 593 | 1.66 | 220 | 2.89 | 2.98 |
| Bayes Flat |  |  |  |  |  |  |  |  |  |  |  |
| Buyes Poiseon num | 1.82 | 4.46 |  |  |  |  |  | 1.66 | 220 |  |  |
| Buyes Gause num | 1.94 | 455 | 271 | 6.4044 | 3.23 | 5.02 | 593 | 1.88 | 290 | 3.44 | 308 |
| Square Root 3/8 | 1.98 | 422 | 2.66 | 6.4032 | 3.23 | 5.01 | 536 | 1.93 | 239 | 3.07 | 300 |
| Fraser-Feid 1 | 2.14 | 487 | 2.84 |  |  |  | 6.44 | 207 | 895 | 5.76 | 310 |
| Fraser-Feid 2 | 1.11 | 196 | 1.05 |  | 0.98 | 5.02 | 2.511 | 120 |  |  |  |
| Fraser-Feid 3 | 2.34 | 5.02 | 310 |  |  |  | 6.48 | 236 |  |  |  |
| Fraser-Feid 4 |  |  |  |  |  |  |  |  |  |  |  |
| Monte Carlo |  |  |  |  |  |  |  |  |  |  |  |
|  |  |  |  |  |  |  |  |  |  |  |  |
| 3 - 4 sigma high | torrect |  |  |  |  |  |  |  |  |  |  |
| 5.4 sigma low | nearty corr |  |  |  |  |  |  |  |  |  |  |
| conrect approximation |  |  |  |  |  |  |  |  |  |  |  |
| poor approximation |  |  |  |  |  |  |  |  |  |  |  |
| input walue |  |  |  |  |  |  |  |  |  |  |  |

## What is a Good (Z) Variable?

Standard Method of MC Testing a Variable:

- "self-test": compare Z with distribution of statistic for MC assuming background only
- i.e. convert back from Z to probability
- Good if $\operatorname{PrMC}(\mathbf{Z}>\mathbf{Z o})=\operatorname{PrGauss}(\mathbf{Z}>\mathbf{Z o})$
- Intuition: want fast convergence to Gaussian

Why not just compare with "right answer"?

- Variables all supposed to give same Z, right?

But it's not really well-defined!

## What is a Bigger Deviation? Part of Significance Definition!

- Measure Non, Noff = (x,y)
- Which values are worse?
- Farther from line $x=\alpha y$ ?
- Angle? Perpendicular?
- Larger $\mathrm{s}=\mathrm{x}-\alpha \mathrm{y}$ ?
- Trying to order 2-dim $\infty$ set!
- Points on (x,y) plane
- Nuisance parameter bites again
- Statistics give different metrics
contours of equal deviation
- Convergence (to Gaussian)?

Which contour?


- Perhaps for large N?
- Enough peaking so overlapping regions dominate integrals?


## Thank you Milagro! Especially Gus and Jordan for making it possible

- I've Learned a Lot (Thanks for explaining!)
- Stimulating Company
- Excellent Surroundings
- Chance to work on some long-deferred things
- Interesting Experiment
- Hope I've contributed something usefull.
- I also hope to find a way to continue...

$$
\begin{gathered}
\text { A Prickly Problem } \\
\text { not to everyone's taste... }
\end{gathered}
$$

- What is Significance?
- Li and Ma Equations
- Frequentist Methods
- Bayesian Methods
- What is Significance, Really?
- To Do


## Conclusions

- Bad formula typically overestimates significance
- For the Crab, any formula will do
- Not true for GRB's with smaller Non, Noff
- LR quite good, though maybe Binomial better
- Several interesting relationships among methods
- Fraser Reid remarkably good for $\mathrm{P}(\mathrm{n} \mid \mathrm{b})$
- Haven't deciphered for interesting case (Non,Noff)
- Binomial Test should be used more


## To Do

- Finish algebra for comparison of Bayes Gamma and Frequentist Binomial
- Monte Carlo Tests
- Fraser Reid Approximation to full problem
- Simpler numerics, if it works!

