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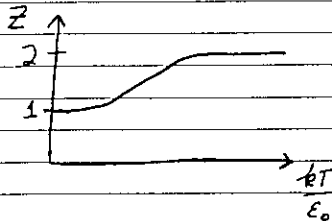
① RB Chapter 5, problem 1

(a) The partition function is

$$Z = \sum_r e^{-E_r/kT}$$

In this problem, there are two energy levels: 0 and ϵ_0 . Thus,

$$Z = 1 + e^{-\epsilon_0/kT}$$



(b) The mean value of the energy is

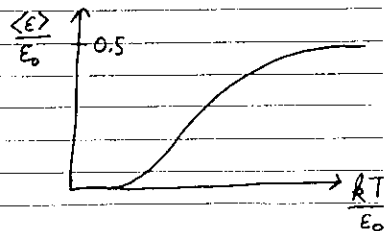
$$\langle E \rangle = \frac{\sum_r E_r e^{-E_r/kT}}{Z}$$

Thus,

$$\langle E \rangle = \frac{\epsilon_0 e^{-\epsilon_0/kT}}{1 + e^{-\epsilon_0/kT}}$$

Multiplying the numerator and denominator by $e^{\epsilon_0/kT}$, we can rewrite $\langle E \rangle$ as follows

$$\langle E \rangle = \frac{\epsilon_0}{e^{\epsilon_0/kT} + 1}$$



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(c) The heat capacity is

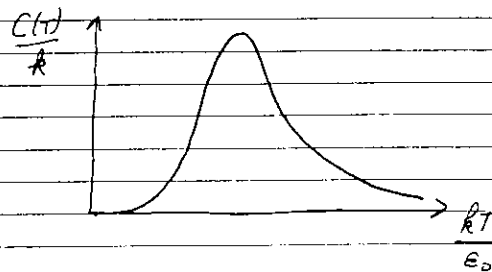
$$C(T) = \frac{\partial \langle E \rangle}{\partial T} = k \left(\frac{\epsilon_0}{kT} \right)^2 \frac{e^{\epsilon_0/kT}}{(e^{\epsilon_0/kT} + 1)^2}$$

To graph this function, we note that as $T \rightarrow 0$, $e^{\epsilon_0/kT} \gg 1$. Thus, we can drop the factor of +1 in the denominator, obtaining:

$$C(T) \approx k \left(\frac{\epsilon_0}{kT} \right)^2 e^{-\epsilon_0/kT}, \quad T \rightarrow 0$$

which is exponentially suppressed. As $T \rightarrow \infty$, $e^{\epsilon_0/kT} = 1 + o\left(\frac{\epsilon_0}{kT}\right)$. Thus,

$$C(T) \approx \frac{k}{4} \left(\frac{\epsilon_0}{kT} \right)^2, \quad T \rightarrow \infty$$

which vanishes like a power. Clearly, $C(T) > 0$ for all T , so $C(T)$ must reach a maximum somewhere in the middle. The graph then looks like

(d) To compute the entropy, use the relation

$$S = \frac{\langle E \rangle}{T} + k \ln Z$$

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Using the results of parts (a) and (b),

$$S = \frac{E_0}{T} \frac{1}{e^{E_0/RT} + 1} + k \ln(1 + e^{-E_0/RT})$$

Look at the cases of small and large T .
As $T \rightarrow 0$,

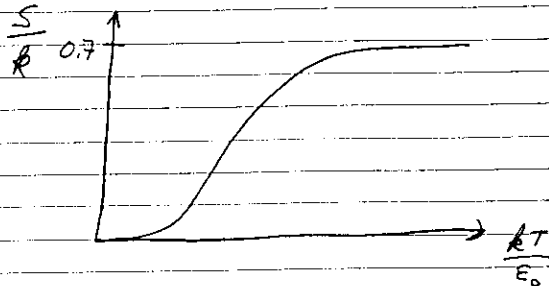
$$\ln(1 + e^{-E_0/RT}) \approx e^{-E_0/RT}$$

since $\ln(1+z) \approx z$ for $|z| \ll 1$. Similarly, we can drop the $+1$ in the denominator of the first term. Thus,

$$S \approx k \left(\frac{E_0}{kT} + 1 \right) e^{-E_0/RT}$$

$$\approx \frac{E_0}{T} e^{-E_0/RT} \quad \text{as } T \rightarrow 0$$

As $T \rightarrow \infty$, $S \rightarrow k \ln 2$. Thus, the graph looks like



noting that $\ln 2 = 0.6931$

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② RB Chapter 5, problem 2

The three energy levels are:

	energy	S_z
spin "up"	E_0	k
spin "sideways"	0	0
spin "down"	E_0	$-k$

Thus, the partition function is

$$Z = 1 + 2e^{-E_0/RT}$$

since the level with energy E_0 is doubly degenerate.

$$(a) P_{up} = \frac{e^{-E_0/RT}}{Z} = \frac{e^{-E_0/RT}}{1 + 2e^{-E_0/RT}}$$

which we can rewrite as

$$P_{up} = \frac{1}{2 + e^{E_0/RT}}$$

In order that $P_{up} = \frac{1}{3}$, we must have $e^{E_0/RT} = 1$ or $\frac{E_0}{RT} = 0$.

This corresponds to the limit of $T \rightarrow \infty$.

(b) Following the same steps as in the previous problem,

$$\langle E \rangle = \frac{1}{Z} 2E_0 e^{-E_0/RT} = 2E_0 P_{up}$$

which makes sense since $P_{up} = P_{down}$. From part (a),

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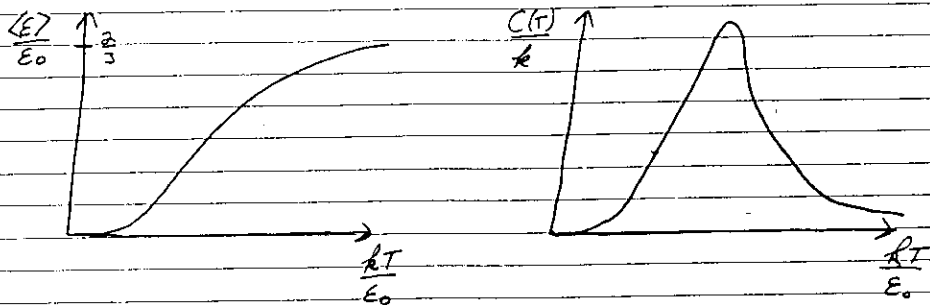
$$\langle E \rangle = \frac{2E_0}{2 + e^{E_0/kT}}$$

The heat capacity is

$$C(T) = \frac{d\langle E \rangle}{dT} = 2k \left(\frac{E_0}{kT} \right)^2 \frac{e^{E_0/kT}}{(2 + e^{E_0/kT})^2}$$

The graphs are similar to those of problem 2. Note that as $T \rightarrow \infty$,

$$\frac{\langle E \rangle}{E_0} \rightarrow \frac{2}{3} \quad \text{Thus,}$$



(c) We can easily work out:

$$\langle S_z \rangle = (+\hbar) P_{up} + (0) P_{sideways} + (-\hbar) P_{down}$$

But $P_{up} = P_{down}$. Thus,

$$\langle S_z \rangle = 0$$

This is to be expected since spin up and spin down occur with equal probability and thus average out to zero.

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3) RB Chapter 5, problem 5

(a) First note that for the case of $N=1$, the energy levels are:

state index j	orientation of moment relative to \vec{B}	system energy E_j
1	parallel	$-m_0 B$
2	anti-parallel	$+m_0 B$

as shown in Table 5.1 of RB (p95). Thus,

$$Z_1 = e^{m_0 B/kT} + e^{-m_0 B/kT} = 2 \cosh \left(\frac{m_0 B}{kT} \right)$$

as obtained in eq. (5.14).

In the case of $N=2$ we have four possible states:

state index j	$E(1)$	$E(2)$	E_j
1	$-m_0 B$	$-m_0 B$	$-2m_0 B$
2	$+m_0 B$	$-m_0 B$	0
3	$-m_0 B$	$+m_0 B$	0
4	$+m_0 B$	$+m_0 B$	$+2m_0 B$

where $E_j = E(1) + E(2)$ is the sum of the single particle energies.

Hence,

$$\begin{aligned} Z_2 &= \sum_{j=1}^4 e^{-E_j/kT} = e^{2m_0 B/kT} + 2 + e^{-2m_0 B/kT} \\ &= (e^{m_0 B/kT} + e^{-m_0 B/kT})^2 \\ &= Z_1^2 \end{aligned}$$

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(b) For the general case of N spins,

$$Z_N = Z_1^N$$

$$\text{Thus, } \ln Z_N = N \ln Z_1$$

We can compute $\langle E \rangle$ from the formula [RB eq (5.16)]

$$\begin{aligned} \langle E \rangle &= kT^2 \frac{1}{Z_N} \frac{\partial Z_N}{\partial T} = kT^2 \frac{\partial}{\partial T} \ln Z_N \\ &= kT^2 N \frac{\partial}{\partial T} \ln Z_1 \\ &= kT^2 N \frac{1}{Z_1} \frac{\partial Z_1}{\partial T} \end{aligned}$$

$$\text{Using } Z_1 = 2 \cosh\left(\frac{m_0 B}{kT}\right),$$

$$\langle E \rangle = -N m_0 B \tanh\left(\frac{m_0 B}{kT}\right)$$

Likewise, the mean value of the magnetic moment along \vec{B} is given by [RB eq 5.22]

$$\begin{aligned} \langle \text{magnetic moment along } \vec{B} \rangle &= kT \frac{\partial}{\partial B} \ln Z_N \\ &= N kT \frac{\partial}{\partial B} \ln Z_1 \\ &= N kT \frac{1}{Z_1} \frac{\partial Z_1}{\partial B} \end{aligned}$$

Evaluating the derivative, one gets

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$$\langle \text{total magnetic moment along } \vec{B} \rangle = N m_0 \tanh\left(\frac{m_0 B}{kT}\right)$$

(c) Consider

$$\langle E \rangle = \frac{1}{Z} \sum_1 E_j e^{-E_j/kT}$$

$$\frac{\partial \langle E \rangle}{\partial T} = \frac{\partial}{\partial T} \left(\frac{1}{Z} \right) \sum_1 E_j e^{-E_j/kT} + \frac{1}{Z} \sum_1 E_j \frac{E_j}{kT^2} e^{-E_j/kT}$$

$$\text{Using } \frac{\partial}{\partial T} \left(\frac{1}{Z} \right) = -\frac{1}{Z^2} \frac{\partial Z}{\partial T} = -\frac{1}{Z} \frac{\langle E \rangle}{kT^2}$$

where the last step follows from RB eq (5.16),

$$\frac{\partial \langle E \rangle}{\partial T} = -\frac{1}{kT^2} \langle E \rangle^2 + \frac{1}{kT^2} \langle E^2 \rangle$$

$$\text{Since } \langle E^2 \rangle = \frac{1}{Z} \sum_1 E_j^2 e^{-E_j/kT}$$

Finally, recall that

$$\begin{aligned} \langle (E - \langle E \rangle)^2 \rangle &= \langle E^2 \rangle - 2 \langle E \langle E \rangle \rangle + \langle E \rangle^2 \\ &= \langle E^2 \rangle - 2 \langle E \rangle^2 + \langle E \rangle^2 \\ &= \langle E^2 \rangle - \langle E \rangle^2 \end{aligned}$$

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Thus,

$$\langle (E - \langle E \rangle)^2 \rangle = kT^2 \frac{\partial \langle E \rangle}{\partial T}$$

For the spin system,

$$\langle E \rangle = -Nm_B \tanh\left(\frac{m_B B}{kT}\right)$$

$$\frac{\partial \langle E \rangle}{\partial T} = \frac{N(m_B B)^2}{kT^2} \operatorname{sech}^2\left(\frac{m_B B}{kT}\right)$$

Using the above formula for $(\Delta E)^2 \equiv \langle (E - \langle E \rangle)^2 \rangle$,

$$\Delta E \equiv \sqrt{N} m_B B \operatorname{sech}\left(\frac{m_B B}{kT}\right)$$

and so,

$$\frac{\Delta E}{|\langle E \rangle|} = \frac{1}{\sqrt{N}} \frac{1}{\sinh\left(\frac{m_B B}{kT}\right)}$$

In particular, we note that provided $\sinh\left(\frac{m_B B}{kT}\right)$ is of order unity,

$$\frac{\Delta E}{|\langle E \rangle|} = O\left(\frac{1}{\sqrt{N}}\right)$$

(d) When $N \sim O(N_A)$ where $N_A \equiv 6.022 \times 10^{23}$ is Avogadro's number, we see that the fractional uncertainty in the energy is incredibly small. A similar computation would also show that the fractional uncertainty in the total magnetic moment along B is also of $O\left(\frac{1}{\sqrt{N}}\right)$ and is likewise negligible.

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(e) The specific heat is given by

$$C(T) \equiv \frac{\partial \langle E \rangle}{\partial T} = \frac{1}{kT^2} (\Delta E)^2$$

Using the results of part (c). This is a general result for the canonical probability distribution in general.

Clearly, $(\Delta E)^2 \geq 0$. Hence we conclude that

$$C(T) \geq 0.$$

Note that one expects that $E \neq \langle E \rangle$ (there is always some thermal fluctuation), in which case $(\Delta E)^2$ is strictly positive. Then, we can conclude that $C(T) > 0$.

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④ RB Chapter 5, problem 7

(a) The partition function is given by [RB eq. (5.22)]

$$Z = \int_0^{\infty} e^{-E/kT} D(E) dE$$

Assuming that $D(E) = g(N, V) E^{\frac{3N}{2}-1}$,

$$Z = g(N, V) \int_0^{\infty} e^{-E/kT} E^{\frac{3N}{2}-1} dE$$

Let $x = \frac{E}{kT}$. Then,

$$Z = g(N, V) (kT)^{\frac{3N}{2}} \int_0^{\infty} e^{-x} x^{\frac{3N}{2}-1} dx$$

$$= g(N, V) (kT)^{\frac{3N}{2}} \Gamma\left(\frac{3N}{2}\right)$$

where Γ is the well-known gamma function. If N is even, we can use $\Gamma(n+1) = n!$ to write:

$$Z = g(N, V) (kT)^{\frac{3N}{2}} \left(\frac{3N}{2} - 1\right)!$$

(b) According to RB eq. (5.24), we can write

$$Z = e^{-\langle E \rangle / kT} D(\langle E \rangle) \delta E$$

which defines δE . To evaluate this, we need to work out $\langle E \rangle$.

(12)

$$\langle E \rangle = \frac{1}{Z} \int_0^{\infty} E e^{-E/kT} D(E) dE$$

$$= \frac{g(N, V)}{Z} \int_0^{\infty} e^{-E/kT} E^{\frac{3N}{2}} dE$$

$$= \frac{g(N, V) (kT)^{\frac{3N}{2}+1} \left(\frac{3N}{2}\right)!}{Z}$$

Inserting $Z = g(N, V) (kT)^{\frac{3N}{2}} \left(\frac{3N}{2} - 1\right)!$ from part (a) yields

$$\langle E \rangle = \frac{3NkT}{2}$$

We could have also derived the same result from

$$\langle E \rangle = kT^2 \frac{\partial}{\partial T} \ln Z$$

Thus,

$$\delta E = \frac{Z}{e^{-\langle E \rangle / kT} D(\langle E \rangle)}$$

$$= \frac{g(N, V) (kT)^{\frac{3N}{2}} \left(\frac{3N}{2} - 1\right)!}{e^{-\frac{3N}{2} kT} g(N, V) \left(\frac{3N}{2} kT\right)^{\frac{3N}{2}-1}}$$

$$= \frac{\left(\frac{3N}{2} - 1\right)! kT}{e^{-\frac{3N}{2} kT} \left(\frac{3N}{2} kT\right)^{\frac{3N}{2}-1}}$$

$$= \frac{\left(\frac{3N}{2} - 1\right)! kT}{e^{-\frac{3N}{2} kT} \left(\frac{3N}{2} kT\right)^{\frac{3N}{2}-1}}$$

$$= \frac{\left(\frac{3N}{2} - 1\right)! kT}{e^{-\frac{3N}{2} kT} \left(\frac{3N}{2} kT\right)^{\frac{3N}{2}-1}}$$

(13)

which can be rewritten as

$$\delta E = \frac{\left(\frac{3N}{2}\right)! kT}{e^{-\frac{3}{2}N} \left(\frac{3N}{2}\right)^{\frac{3N}{2}}}$$

Finally, using Stirling's approximation,

$$n! \approx \sqrt{2\pi n} \left(\frac{n}{e}\right)^n, \quad n \gg 1$$

we can write:

$$\delta E \approx \frac{\sqrt{2\pi} \left(\frac{3N}{2}\right)^{\frac{3N}{2} + \frac{1}{2}} e^{-\frac{3}{2}N} kT}{e^{-\frac{3}{2}N} \left(\frac{3N}{2}\right)^{\frac{3N}{2}}}$$

or

$$\delta E \approx \sqrt{3\pi N} kT$$

(c) In part (b), we have already worked out $\langle E \rangle = \frac{2}{3} N kT$.
Thus,

$$\frac{\delta E}{\langle E \rangle} = \frac{2}{3} \sqrt{\frac{3\pi}{N}}$$

In particular, we note that $\frac{\delta E}{\langle E \rangle} = O\left(\frac{1}{\sqrt{N}}\right)$

and again we find that the fractional uncertainty in the energy is completely negligible when N is a macroscopic number, say $N = O(N_A)$.